## The Kolmogorov Law of turbulence

 What can rigorously be proved ?Roger LEWANDOWSKI

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## Introduction



Aim: Mathematical framework for the Kolomogorov laws.

## Table of contents

(1) Incompressible Navier-Stokes Equations (NSE), (2) Probabilistic framework

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(1) Incompressible Navier-Stokes Equations (NSE),
(2) Probabilistic framework, Reynolds Stress, Correlations, (3) Homogeneity,

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(1) Incompressible Navier-Stokes Equations (NSE),
(2) Probabilistic framework, Reynolds Stress, Correlations,
(3) Homogeneity, Turbulent Kinetic Energy, Dissipation,
(4) Isotropy, Energy spectrum, Similarity and law of the $-5 / 3$.

## Introduction



Authors in Maths publications are always by alphabetical order

## The Foundations of Chaos Revisited: From Poincaré to Recent Advancements

Lewandowski, R and Pinier, B., The Kolmogorov-Taylor Law of turbulence : what can rigorously be proved ? Part II, In : The Foundations of Chaos Revisited: From Poincaré to Recent Advancements, 82-101, Springer, 2016.

## Solutions to the NSE

1) Incompressible 3D Navier-Stokes Equations (NSE)

$$
Q=[0, T] \times \Omega \quad \text { or } \quad Q=\left[0, \infty\left[\times \Omega, \quad \Omega \subset \mathbb{R}^{3}, \quad \Gamma=\partial \Omega,\right.\right.
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with the no slip boundary condition and $v_{0}$ as initial data:

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D \mathbf{v}=\frac{1}{2}\left(\nabla \mathbf{v}+\nabla \mathbf{v}^{t}\right), \quad((\mathbf{v} \cdot \nabla) \mathbf{v})_{i}=v_{j} \frac{\partial v_{i}}{\partial x_{j}}, \quad \nabla \cdot \mathbf{v}=\frac{\partial v_{i}}{\partial x_{i}} .
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## Remark

$\nabla \cdot \mathbf{v}=0 \Rightarrow(\mathbf{v} \cdot \nabla) \mathbf{v}=\nabla \cdot(\mathbf{v} \otimes \mathbf{v}), \quad \mathbf{v} \otimes \mathbf{v}=\left(v_{i} v_{j}\right)_{1 \leq i, j \leq 3}$.

## Solutions to the 3D NSE

2) Two types of solutions to the 3D NSE
(1) Strong solutions over a small time interval $\left[0, T_{\max }[\right.$ "à la" Fujita-Kato,
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\mathbf{v}=\mathbf{v}\left(t, \mathbf{x}, \mathbf{v}_{0}\right), \quad p=p\left(t, \mathbf{x}, \mathbf{v}_{0}\right)
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## Remark

Strong solutions are defined over $\left[0, \infty\left[\right.\right.$ when $\mathrm{v}_{0}$ and /or ||f|| are
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Weak solutions are defined through an appropriate variational formulation set in the sequence of function spaces
$\left\{\mathbf{v} \in H_{0}^{1}(\Omega)^{3}, \nabla \cdot \mathbf{v}=0\right\} \hookrightarrow \mathbf{V}=\left\{\mathbf{v} \in L^{2}(\Omega)^{3},\left.\mathbf{v} \cdot \mathbf{n}\right|_{\Gamma}=0, \nabla \cdot \mathbf{v}=0\right\}$
such that the trajectory
is weakly continuous from $[0, T] \rightarrow \mathbf{V}, T \in] 0, \infty](\forall \boldsymbol{\eta} \in \mathbf{V}$, $t \rightarrow\langle\mathbf{v}(t), \boldsymbol{\eta}\rangle$ is a continuous function of $t$, where $\langle\cdot, \cdot\rangle$ denotes the scalar product in V ).

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## Remark

For $\mathbf{v}_{0} \in \mathbf{V}$ and $\mathbf{f} \in L^{2}(\Omega)^{3}$, there exists a global weak solution $\mathbf{v}=\mathbf{v}(t), t \in \mathbb{R}^{+}$. Because of lack of uniqueness result, we can't write

$$
\mathbf{v}=\mathbf{v}\left(t, \mathbf{x}, \mathbf{v}_{0}\right)
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## Probabilistic framework

## 1) Long Time Average

Let $\mathcal{B}\left(\mathbb{R}_{+}\right)$denotes the Borel $\sigma$-algebra on $\mathbb{R}_{+}, \lambda$ the Lebesgue measure, and let $\mu$ denotes the "probability half-measure"

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\forall A \in \mathcal{B}\left(\mathbb{R}_{+}\right), \quad \mu(A)=\lim _{t \rightarrow \infty} \frac{1}{t} \lambda(A \cap[0, t]),
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\text { Let } \mathbf{v} \in L^{1}\left(\mathbb{R}^{+} \rightarrow \mathbf{V} ; \mu\right)
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E(\mathbf{v})=\overline{\mathbf{v}}=\int_{\mathbf{R}^{+}} \mathbf{v}(t) d \mu(t)=\lim _{t \rightarrow \infty} \frac{1}{t} \int_{0}^{t} \mathbf{v}(t) d t
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Let $v$ be a Leray-Hopf solution to the NSE. It is not known wether for a given $x$
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Let $\mathbf{v}$ be a Leray-Hopf solution to the NSE. It is not known wether for a given $\mathbf{x} \in \Omega, \mathbf{v}(t, \mathbf{x}) \in L^{1}\left(\mathbb{R}^{+} \rightarrow \mathbf{V} ; \mu\right)$. However, it can be proved that in some sense when $\mathbf{f}$ is steady ${ }^{1}$,

$$
E(\mathbf{v})=\overline{\mathbf{v}}=\overline{\mathbf{v}}(\mathbf{x}) \in W^{2,5 / 4}(\Omega)^{3}
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[^0]
## Probabilistic framework

## 2) Ensemble Average

The source term $\mathbf{f}$ and the viscosity $\nu$ are fixed. Let $\mathbb{K} \subset \mathbf{V}$ be a compact,

$$
T_{\mathbb{K}}=\inf _{\mathbf{v}_{0} \in \mathbb{K}} T_{\max }\left(\left\|\mathbf{v}_{0}\right\|,\|\mathbf{f}\|, \nu\right)>0, \quad Q=\left[0, T_{\mathbb{K}}\right] \times \Omega
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where $\mu_{n}$ is the probability measure over $\mathbb{K}$ :

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Let $\left\{\mathbf{v}_{0}^{(1)}, \cdots, \mathbf{v}_{0}^{(n)}, \cdots\right\}$ be a countable dense subset of $\mathbb{K}$,

$$
\overline{\mathbf{v}}_{n}=\overline{\mathbf{v}}_{n}(t, \mathbf{x})=\frac{1}{n} \sum_{i=1}^{n} \mathbf{v}\left(t, \mathbf{x}, \mathbf{v}_{0}^{(i)}\right)=E_{\mu_{n}}(\mathbf{v}(t, \mathbf{x}, \cdot))
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## Remark

It is not known if the probability measure $\mu$ is unique or not

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## Reynolds Stress

1) Reynolds decomposition We can decompose ( $\mathbf{v}, p$ ) as follows:

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\mathbf{v}=\overline{\mathbf{v}}+\mathbf{v}^{\prime}, \quad p=\bar{p}+p^{\prime}
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which is the Reynolds decomposition, $\mathbf{v}^{\prime}$ and $p^{\prime}$ are the fluctuations.

Either for long time or ensemble averages:

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## Lemma

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Lemma

$$
\forall(t, \mathbf{x}) \in Q, \quad \overline{\mathbf{v}^{\prime}\left(t, \mathbf{x}, \mathbf{v}_{0}\right)}=0, \quad \overline{p^{\prime}\left(t, \mathbf{x}, \mathbf{v}_{0}\right)}=0
$$

## Reynolds Stress

## 2) Averaged NSE

Note that $E_{\mu}(\mathbf{f})=\mathbf{f} E_{\mu}(1)=\mathbf{f}$. By the Reynolds rules and the previous lemma:

$$
\left\{\begin{aligned}
\partial_{t} \overline{\mathbf{v}}+(\overline{\mathbf{v}} \cdot \nabla) \overline{\mathbf{v}}-\nu \Delta \overline{\mathbf{v}}+\nabla \bar{p} & =-\nabla \cdot \boldsymbol{\sigma}^{(\mathrm{R})}+\mathbf{f} & & \text { in } Q \\
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where

$$
\boldsymbol{\sigma}^{(\mathrm{R})}=\overline{\mathbf{v}^{\prime} \otimes \mathbf{v}^{\prime}}
$$

is the Reynolds stress.

## Standard correlation tensors

Standard correlation tensor Let $M_{1}, \ldots, M_{n} \in Q, M_{k}=\left(t_{k}, \mathbf{x}_{k}\right)$,

$$
\mathbb{B}_{n}=\mathbb{B}_{n}\left(M_{1}, \ldots, M_{n}\right)=\left(B_{i_{1} \ldots i_{n}}\left(M_{1}, \ldots, M_{n}\right)\right)_{1 \leq i_{1} \ldots i_{n} \leq 3}
$$

at these points is defined by

$$
\begin{aligned}
& B_{i_{1} \ldots i_{n}}\left(M_{1}, \ldots, M_{n}\right)=\prod_{k=1}^{n} v_{i_{k}}\left(t_{k}, \mathbf{x}_{k}, \mathbf{v}_{0}\right)
\end{aligned}=\begin{aligned}
& \int_{\mathbb{K}}\left(\prod_{k=1}^{n} v_{i_{k}}\left(t_{k}, \mathbf{x}_{k}, \mathbf{v}_{0}\right)\right) d \mu\left(\mathbf{v}_{0}\right)=E_{\mu}\left(\prod_{k=1}^{n} v_{i_{k}}\left(t_{k}, \mathbf{x}_{k}, \mathbf{v}_{0}\right)\right)
\end{aligned}
$$

where $\mathbf{v}=\left(v_{1}, v_{2}, v_{3}\right)$.

## Discussion about Homogeneity

## Extension of the test family

(1) field family:

$$
\mathcal{G}=\left\{\begin{array}{c}
\left.v_{1}, v_{2}, v_{3}, p, \partial_{j} v_{i}(1 \leq i, j \leq 3)\right\}, \partial_{t} v_{i}(1 \leq i \leq 3) \\
\partial_{i} p(1 \leq i \leq 3), \partial_{i j}^{2} v_{k}(1 \leq i, j, k \leq 3)
\end{array}\right\}
$$

(2) fluctuations field family:

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\partial_{i} p(1 \leq i \leq 3), \partial_{i j}^{2} v_{k}(1 \leq i, j, k \leq 3)
\end{array}\right\}
$$

(2) fluctuations field family:

$$
\mathcal{H}=\left\{\begin{array}{c}
\left.v_{1}^{\prime}, v_{2}^{\prime}, v_{3}^{\prime}, p^{\prime}, \partial_{j} v_{i}^{\prime}(1 \leq i, j \leq 3)\right\}, \partial_{t} v_{i}^{\prime}(1 \leq i \leq 3) \\
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$$

each element of
is Hölder continuous with respect to $(t, x) \in Q$, and continuous

## Discussion about Homogeneity

## Extension of the test family

(1) field family:

$$
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\end{array}\right\}
$$

each element of

$$
\mathcal{F}=\mathcal{G} \cup \mathcal{H}
$$

is Hölder continuous with respect to $(t, \mathbf{x}) \in Q$, and continuous with respect to $\mathbf{v}_{0} \in \mathbb{K}$.

## Discussion about Homogeneity

Let $D=I \times \omega \subset Q$ open and connected subset, such that $l \subset] 0, T_{\mathbb{K}}[$ and $\bar{\omega} \subset \Omega$.

Aim To introduce different concepts of homogeneity in $D$, reflected in the local invariance under spatial translations of the correlation tensors based on the families $\mathcal{G}$ and/or $\mathcal{F}=\mathcal{G} \cup \mathcal{H}$, which is essential in the derivation of models such as $k-\mathscr{E}$.

Let $M=(t, x) \in D$, and denote

For simplicity, we also denote

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Let $M=(t, \mathbf{x}) \in D$, and denote

$$
\left(\tau_{t}, r_{\mathbf{x}}\right)=\sup \{(t, r) /] t-\tau, t+\tau[\times B(\mathbf{x}, r) \subset D\}
$$

For simplicity, we also denote

$$
(t+\tau, \mathbf{x}+\mathbf{r})=M+(\tau, \mathbf{r}), \quad(t, \mathbf{x}+\mathbf{r})=M+\mathbf{r} .
$$

## Discussion about Homogeneity

## Definition

We say that the flow is 1 ) homogeneous (standard definition), 2) strongly homogeneous (extended definition, suitable for $k-\mathscr{E}$ ) in $D$, if $\forall n \in \mathbb{N}$,

$$
\text { 1) } \forall M_{1}, \ldots, M_{n} \in D, \quad \forall \psi_{1}, \ldots, \psi_{n} \in \mathcal{G}, \quad \forall \mathbf{r} \in \mathbb{R}^{3}
$$

$$
\text { 2) } \forall M_{1}, \ldots, M_{n} \in D, \quad \forall \psi_{1}, \ldots, \psi_{n} \in \mathcal{F}, \quad \forall \mathbf{r} \in \mathbb{R}^{3}
$$

such that $|\mathbf{r}| \leq \inf _{1 \leq i \leq n} r_{\mathrm{x}_{\mathrm{i}}}$, we have
$B\left(\psi_{1}, \ldots, \psi_{n}\right)\left(M_{1}+\mathbf{r}, \ldots, M_{n}+\mathbf{r}\right)=B\left(\psi_{1}, \ldots, \psi_{n}\right)\left(M_{1}, \ldots, M_{n}\right)$,
where

$$
B\left(\psi_{1}, \ldots, \psi_{n}\right)\left(M_{1}, \ldots, M_{n}\right)=\overline{\psi_{1}\left(M_{1}\right) \cdots \psi_{n}\left(M_{n}\right)},
$$

## Discussion about Homogeneity

## Lemma

Assume that the flow is homogeneous (resp. strongly hom.). Let

$$
\psi_{1}, \ldots, \psi_{n} \in \mathcal{G}(r e s p \in \mathcal{F}), \quad M_{1}, \ldots, M_{n} \in D, \quad M_{i}=\left(t_{i}, \mathbf{r}_{i}\right),
$$

such that

$$
\forall i=1, \cdots, n, \quad t_{i}=t
$$

Let $\mathbf{r}_{i}$ denotes the vector such that $M_{i}=M_{1}+\mathbf{r}_{i-1}(i \geq 2)$. Then, $B\left(\psi_{1}, \ldots, \psi_{n}\right)\left(M_{1}, \ldots, M_{n}\right)$ only depends on $t$ and $\mathbf{r}_{1}, \cdots, \mathbf{r}_{n-1}$.

## Discussion about Homogeneity

## Theorem

Assume that $\mathbf{f}$ satisfies the compatibility condition $\nabla \mathbf{f}=0$ in $D$, and the flow is strongly homogeneous in D. Then
(1) $\forall \psi \in \mathcal{F}, \nabla \bar{\psi}=0$ in $D$,
(2) $\nabla \sigma^{(\mathrm{R})}=0$ in $D$,
(3) and we have $\forall t \in I$,

$$
\overline{\mathbf{v}}=\overline{\mathbf{v}}(t)=\overline{\mathbf{v}}\left(t_{0}\right)+\int_{t_{0}}^{t} \mathbf{f}(s) d s \text { in } D
$$

by noting $t_{0}=\inf I$.

## Discussion about Homogeneity

## Definition

We say that a flow is mildly homogneous in $D=I \times \omega$, if $\forall \psi, \varphi \in \mathcal{H}$ (fluctuations family), we have

$$
\forall M=(t, \mathbf{x}) \in D, \quad \overline{\psi(t, \mathbf{x}) \partial_{i} \varphi(t, \mathbf{x})}=-\overline{\partial_{i} \psi(t, \mathbf{x}) \varphi(t, \mathbf{x})} .
$$

This definition is motivated by:

## Lemma

Any strongly homogeneous flow is mildly homogeneous.

## Equations for the TKE and the turbulent dissipation

The turbulent kinetic energy $k$ (TKE) and the turbulent dissipation $\mathscr{E}$ are defined by:

$$
k=\frac{1}{2} \operatorname{tr} \sigma^{(\mathrm{R})}=\frac{1}{2} \overline{\left|\mathbf{v}^{\prime}\right|^{2}}, \quad \mathscr{E}=2 \nu \overline{\mid \overline{\left.\mathbf{v}^{\prime}\right|^{2}}} .
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$$

Question: What equation can we find out for $k$ and $\mathscr{E}$ ?

## Equations for the TKE and the turbulent dissipation

## Theorem

Assume that the flow is mildly homogeneous in $D$
(1) The TKE $k$ satisfies in D:

$$
\partial_{t} k+\overline{\mathbf{v}} \cdot \nabla k+\nabla \cdot \overline{e^{\prime} \mathbf{v}^{\prime}}=-\boldsymbol{\sigma}^{(\mathrm{R})}: \nabla \overline{\mathbf{v}}-\mathscr{E}
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(2) The turbulent dissipation $\mathscr{E}$ satisfies in $D$
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$$
\begin{aligned}
\partial_{t} \mathscr{E}+\overline{\mathbf{v}} \cdot \nabla \mathscr{E}+\nabla \cdot \overline{\nu h^{\prime} \mathbf{v}^{\prime}}= & 2 \nu\left(\overline{\boldsymbol{\omega}^{\prime} \otimes \boldsymbol{\omega}^{\prime}}: \nabla \overline{\mathbf{v}}+\overline{\left(\boldsymbol{\omega}^{\prime} \otimes \boldsymbol{\omega}^{\prime}\right)^{\prime}: \nabla \mathbf{v}^{\prime}}\right) \\
& -2 \nu^{2}\left|\nabla \boldsymbol{\omega}^{\prime}\right|^{2},
\end{aligned}
$$

where

$$
\begin{array}{lll}
e=\frac{1}{2}\left|\mathbf{v}^{\prime}\right|^{2}, & \text { decomposed as } & e=\bar{e}+e^{\prime}=k+e^{\prime} \\
\boldsymbol{\omega}=\nabla \times \mathbf{v}, & \text { decomposed as } & \boldsymbol{\omega}=\overline{\boldsymbol{\omega}}+\omega^{\prime} \\
h=\left|\boldsymbol{\omega}^{\prime}\right|^{2}, & \text { decomposed as } & h=\bar{h}+h^{\prime}
\end{array}
$$

## Discussion about isotropy

## 1) Basics

Throughout what follows, we assume that the flow is homogeneous (standard definition), and for simplicity stationnary ("homogeneity in time"). Let $\mathcal{B}_{n}$ denotes all correlation tensors of the form:

$$
\begin{aligned}
& \mathbb{B}_{n}=\mathbb{B}_{n}\left(M_{1}, \ldots, M_{n}\right)=\left(B_{i_{1} \ldots i_{n}}\left(M_{1}, \ldots, M_{n}\right)\right)_{1 \leq i_{1} \ldots i_{n} \leq 3} \\
& \psi_{i_{1}}, \ldots, \psi_{i_{n}} \in \mathcal{G} \\
& B_{i_{1} \cdots i_{n}}\left(\psi_{1}, \ldots, \psi_{n}\right)\left(\mathbf{r}_{1}, \ldots, \mathbf{r}_{n-1}\right)=\overline{\psi_{i_{1}}(\mathbf{x}) \psi_{i_{2}}\left(\mathbf{x}+\mathbf{r}_{1}\right) \cdots \psi_{i_{n}}\left(\mathbf{x}+\mathbf{r}_{n-1}\right)}
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using the Einstein summation convention.

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\end{aligned}
$$

Let $\mathbf{a}_{1}, \cdots, \mathbf{a}_{n} \in \mathbb{R}^{3}, \mathbf{a}_{i}=\left(a_{i 1}, a_{i 2}, a_{i 3}\right)$. We set
$\left[B_{n}\left(\mathbf{r}_{1}, \cdots, \mathbf{r}_{n-1}\right),\left(\mathbf{a}_{1}, \cdots, \mathbf{a}_{n}\right)\right]=a_{1 i_{1}} \cdots a_{n i_{n}} B_{i_{1} \cdots i_{n}}\left(\mathbf{r}_{1}, \cdots, \mathbf{r}_{n-1}\right)$,
using the Einstein summation convention.

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$$

using the Einstein summation convention.
We denote by $\mathrm{O}_{3}(\mathbb{R})$ an orthogonal group, which means that $Q \in O_{3}(\mathbb{R})$ if and only if $Q Q^{t}=Q^{t} Q=\mathrm{I}$.

## Discussion about isotropy

## Definition

We say that the flow is isotropic in $D$ if and only if,

$$
\begin{aligned}
& \forall n \geq 1, \quad \forall \mathbb{B}_{n} \in \mathcal{B}_{n}, \\
& \forall Q \in O_{3}(\mathbb{R}), \quad \forall \mathbf{a}_{1}, \cdots, \mathbf{a}_{n} \in \mathbb{R}^{3}, \\
& \forall \mathbf{x} \in \omega, \quad \forall\left(\mathbf{r}_{1}, \cdots, \mathbf{r}_{n-1}\right) \in B\left(0, r_{\mathbf{x}}\right)^{n-1},
\end{aligned}
$$

then we have

$$
\begin{aligned}
& {\left[\mid B_{n}\left(Q \mathbf{r}_{1}, \cdots, Q \mathbf{r}_{n-1}\right),\left(Q \mathbf{a}_{1}, \cdots, Q \mathbf{a}_{n}\right)\right]=} \\
& {\left[B_{n}\left(\mathbf{r}_{1}, \cdots, \mathbf{r}_{n-1}\right),\left(\mathbf{a}_{1}, \cdots, \mathbf{a}_{n}\right)\right] .}
\end{aligned}
$$

## Discussion about isotropy

## 2) Two order tensor

We fix $\delta_{0}$ once and for all and $\mathbf{x}$ satisfies $d(\mathbf{x}, \partial \omega) \geq \delta_{0}$ so that $\mathbb{B}_{2}(\mathbf{r})$ is well defined for $|\mathbf{r}| \leq \delta_{0}$ and at least of class $C^{1}$ with respect to $\mathbf{r}$ (and does not depend on $\mathbf{x}$ ).

## Theorem

Assume the flow isotropic in D. Then there exist two $C^{1}$ scalar functions $B_{d}=B_{d}(r)$ and $B_{n}=B_{n}(r)$ on $\left[0, \delta_{0}[\right.$ and such that

$$
\forall \mathbf{r} \in B\left(0, \delta_{0}\right), \quad \mathbb{B}_{2}(\mathbf{r})=\left(B_{d}(r)-B_{n}(r)\right) \frac{\mathbf{r} \otimes \mathbf{r}}{r^{2}}+B_{n}(r) I_{3}
$$

where $r=|\mathbf{r}|, \mathbf{r} \otimes \mathbf{r}=\left(r_{i} r_{j}\right)_{1 \leq i, j \leq 3}$. Moreover, $B_{d}$ and $B_{n}$ are linked through the following differential relation.

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$$
\forall r \in\left[0, \delta_{0}\left[, \quad r B_{d}^{\prime}(r)+2\left(B_{d}(r)-B_{n}(r)\right)=0\right.\right.
$$

where $B_{d}^{\prime}(r)$ is the derivative of $B_{d}$.

## Energy spectrum

Energy spectrum for isotropic flows Let

$$
\mathbb{E}=\left.\frac{1}{2} \operatorname{tr} \mathbb{B}_{2}\right|_{\mathbf{r}=0}=\frac{1}{2} \overline{|\mathbf{v}|^{2}},
$$

be the total mean kinetic energy

## Theorem <br> There exists a measurable function $E=E(k)$, defined over $\mathbb{R}_{-}$ the integral of which over $\mathbb{R}_{+}$is finite, and such that

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There exists a measurable function $E=E(k)$, defined over $\mathbb{R}_{+}$, the integral of which over $\mathbb{R}_{+}$is finite, and such that

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\mathbb{E}=\int_{0}^{\infty} E(k) d k
$$

Remark
$E(k)$ is the amount of kinetic energy in the sphere $S_{k}=\{|k|=k$
which physically means $E \geq 0$ in $\mathbb{R}_{+}$, and therefore $E \in L^{1}\left(\mathbb{R}_{+}\right)$. We cannot rigorously prove $E \geq 0$ which remains an open problem.

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## Energy spectrum

## Lemma

The turbulent dissipation $\mathscr{E}$ is deduced from the energy spectrum from the formula:

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\mathscr{E}=\nu \int_{0}^{\infty} k^{2} E(k) d k,
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which also states that when $E \geq 0$, then $k^{2} E(k) \in L^{1}\left(\mathbb{R}_{+}\right)$.

The issue is the determination of the profil of $E$

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## Similarity

## 1) Dimensional bases

Only length and time are involved in this frame, heat being not considered and the fluid being incompressible.

## Definition

A length-time basis is a couple $b=(\lambda, \tau)$, where $\lambda$ a given constant length and $\tau$ a constant time.

## Definition

Let $\psi=\psi(t, x)$ (constant, scalar, vector, tensor...) be defined on
 $\Omega$. The couple $\left(d_{\ell}(\psi), d_{\tau}(\psi)\right) \in \mathbb{Q}^{2}$ is such that

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Let $\psi=\psi(t, \mathbf{x})$ (constant, scalar, vector, tensor...) be defined on $Q=\left[0, T_{\mathbb{K}}\right] \times \Omega$. The couple $\left(d_{\ell}(\psi), d_{\tau}(\psi)\right) \in \mathbb{Q}^{2}$ is such that

$$
\psi_{b}\left(t^{\prime}, \mathbf{x}^{\prime}\right)=\lambda^{-d_{\ell}(\psi)} \tau^{-d_{\tau}(\psi)} \psi\left(\tau t^{\prime}, \lambda \mathbf{x}^{\prime}\right)
$$

where $\left(t^{\prime}, \mathbf{x}^{\prime}\right) \in Q_{b}=\left[0, \frac{T_{\mathbb{K}}}{\tau}\right] \times \frac{1}{\lambda} \Omega$, is dimensionless. We say that $\psi_{b}=\psi_{b}\left(t^{\prime}, \mathbf{x}^{\prime}\right)$ is the $b$-dimensionless field deduced from $\psi$.

## Similarity

2) Kolmogorov scales

Let us consider the length-time basis $b_{0}=\left(\lambda_{0}, \tau_{0}\right)$, determined by

$$
\lambda_{0}=\nu^{\frac{3}{4}} \mathscr{E}^{-\frac{1}{4}}, \quad \tau_{0}=\nu^{\frac{1}{2}} \mathscr{E}^{-\frac{1}{2}} .
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We recall that $\lambda_{0}$ is called the Kolmogorov scale.
The important point here is that

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\mathscr{E}_{b_{0}}=\nu_{b_{0}}=1
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Moreover, for all wave number $k$,


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Moreover, for all wave number $k$,

$$
E(k)=\nu^{\frac{5}{4}} \mathscr{E}^{\frac{1}{4}} E_{b_{0}}\left(\lambda_{0} k\right),
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We must find out the universal profil $E_{b_{0}}$

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## Similarity

## 3) Assumptions

Scale separation. Let $\ell$ be the Prandtl mixing length. Then

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Similarity. There exists an interval

$$
\begin{gathered}
{\left[k_{1}, k_{2}\right] \subset\left[\frac{2 \pi}{\ell}, \frac{2 \pi}{\lambda_{0}}\right] \text { s.t. } k_{1} \ll k_{2} \text { and on }\left[\lambda_{0} k_{1}, \lambda_{0} k_{2}\right],} \\
\forall b_{1}=\left(\lambda_{1}, \tau_{1}\right), b_{2}=\left(\lambda_{2}, \tau_{2}\right) \text { s.t. } \mathscr{E}_{b_{1}}=\mathscr{E}_{b_{2}}, \text { then } E_{b_{1}}=E_{b_{2}}
\end{gathered}
$$

## Similarity

## 4) Law of the $-5 / 3$

## Theorem

Scale separation and Similarity Assumptions yield the existence of a constant $C$ such that

$$
\forall k^{\prime} \in\left[\lambda_{0} k_{1}, \lambda_{0} k_{2}\right]=J_{r}, \quad E_{b_{0}}\left(k^{\prime}\right)=C\left(k^{\prime}\right)^{-\frac{5}{3}} .
$$

## Corollary <br> The energy spectrum satisfies the $-5 / 3$ law

where $C$ is a dimensionless constant.

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## Corollary

The energy spectrum satisfies the $-5 / 3$ law

$$
\forall k \in\left[k_{1}, k_{2}\right], \quad E(k)=C \mathscr{E}^{\frac{2}{3}} k^{-\frac{5}{3}},
$$

where $C$ is a dimensionless constant.

## Similarity

## Idea of the proof Let

$$
b^{(\alpha)}=\left(\alpha^{3} \lambda_{0}, \alpha^{2} \tau_{0}\right)
$$

The similarity assumption yields

## which leads to the functional equation,

## Similarity

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$$

As

$$
\mathscr{E}_{b^{(\alpha)}}=1=\mathscr{E}_{b_{0}},
$$

The similarity assumption yields

$$
\forall k^{\prime} \in J_{r}, \quad \forall \alpha>0, \quad E_{b(\alpha)}\left(k^{\prime}\right)=E_{b_{0}}\left(k^{\prime}\right) .
$$

which leads to the functional equation,
whose unique solution is given by

## Similarity

## Idea of the proof Let

$$
b^{(\alpha)}=\left(\alpha^{3} \lambda_{0}, \alpha^{2} \tau_{0}\right)
$$

As

$$
\mathscr{E}_{b^{(\alpha)}}=1=\mathscr{E}_{b_{0}},
$$

The similarity assumption yields

$$
\forall k^{\prime} \in J_{r}, \quad \forall \alpha>0, \quad E_{b^{(\alpha)}}\left(k^{\prime}\right)=E_{b_{0}}\left(k^{\prime}\right)
$$

which leads to the functional equation,

$$
\forall k^{\prime} \in J_{r}, \quad \forall \alpha>0, \quad \frac{1}{\alpha^{5}} E_{b_{0}}\left(k^{\prime}\right)=E_{b_{0}}\left(\alpha^{3} k^{\prime}\right),
$$

whose unique solution is given by

$$
\forall k^{\prime} \in J_{r}, \quad E_{b_{0}}\left(k^{\prime}\right)=C\left(k^{\prime}\right)^{-\frac{5}{3}}, \quad C=\left(\frac{k_{1}}{\lambda_{0}}\right)^{\frac{5}{3}} E_{0}\left(\frac{k_{1}}{\lambda_{0}}\right),
$$

hence the result.

## Consequences

## Theorem

Assume:
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Then Smagorinsky's postulate holds true, i.e

$$
\nu_{t}=C \delta^{2}|D \overline{\mathbf{v}}|
$$


[^0]:    ${ }^{1}$ Chacon-Lewandowski (Springer 2014), Lewandowski (Chin.An.Maths, 2015)

[^1]:    where

